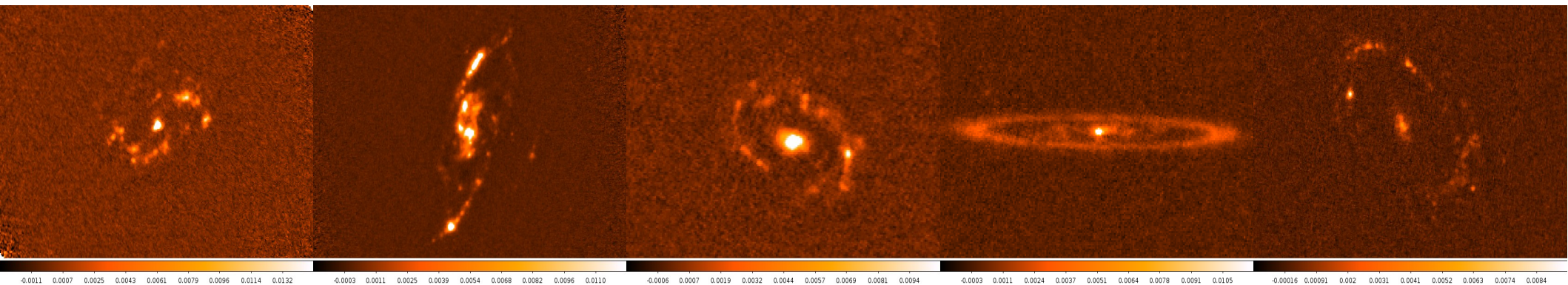


The nuclear and integrated FIR emission of Seyfert galaxies

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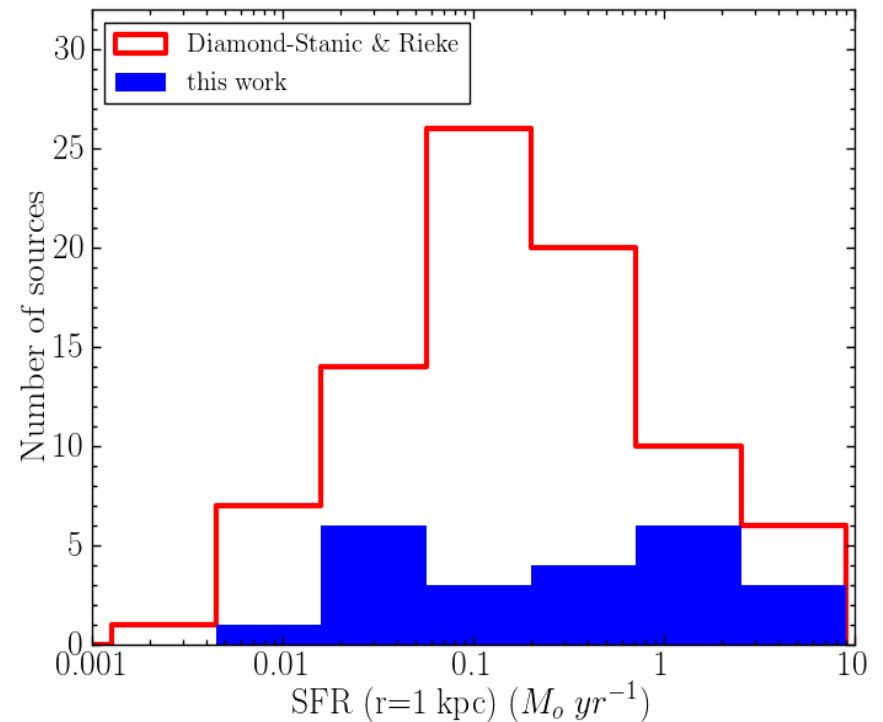
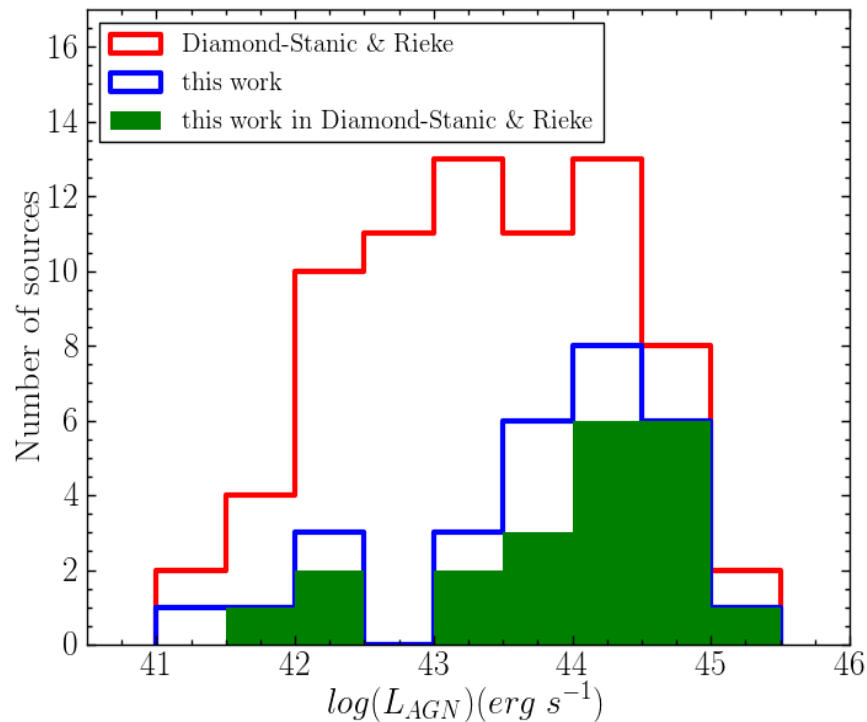


Goals

- Study the FIR nuclear and integrated emission of local Seyfert galaxies taking advantage of Herschel angular resolution.
- Select galaxies with AGN-dominated nuclear 70 μm emission. Four criteria:
 - Elevated 70/160 μm flux ratios with respect to the typical colours of star forming galaxies.
 - Nuclear dust temperature higher than typical values of star forming galaxies.
 - 70 μm excess emission with respect to the fit of the FIR SEDs with a grey body.
 - Comparison of nuclear SFR obtained from 70 μm and mid-IR indicators.

The sample and observations

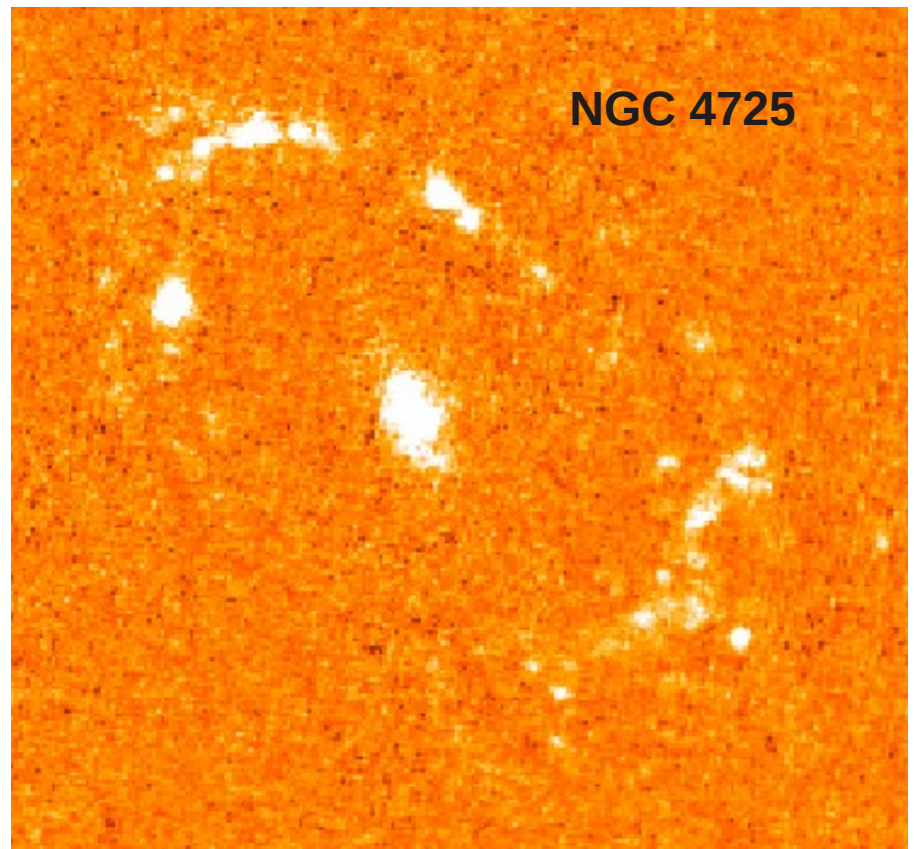
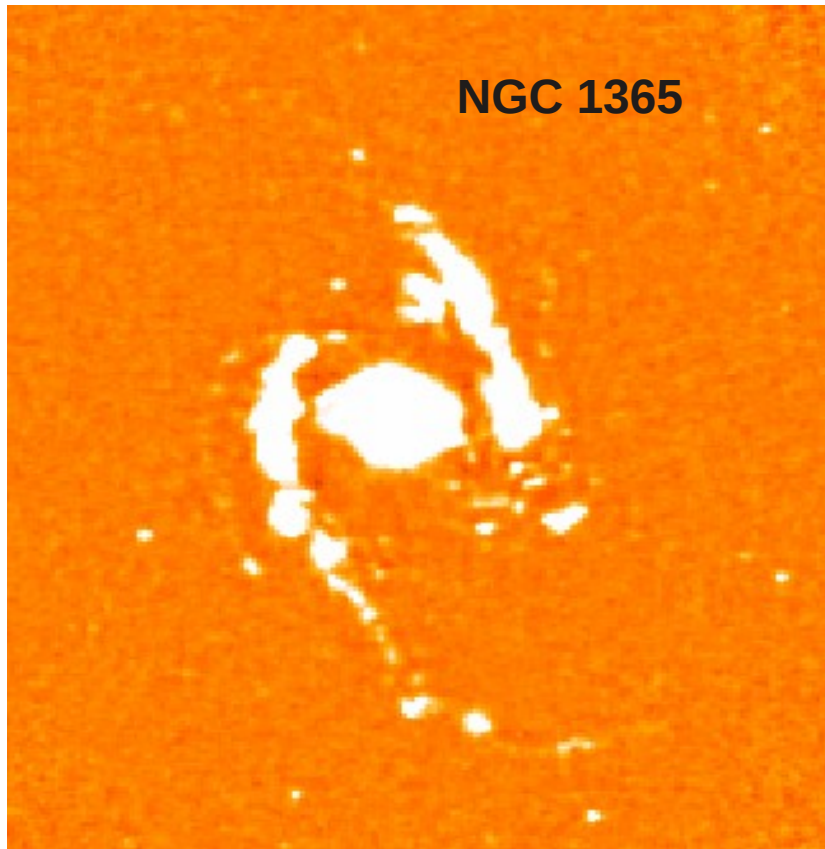
- 33 local Seyfert galaxies from the Revised Shapley-Ames catalogue (RSA sample)



Median distance of 30 Mpc.

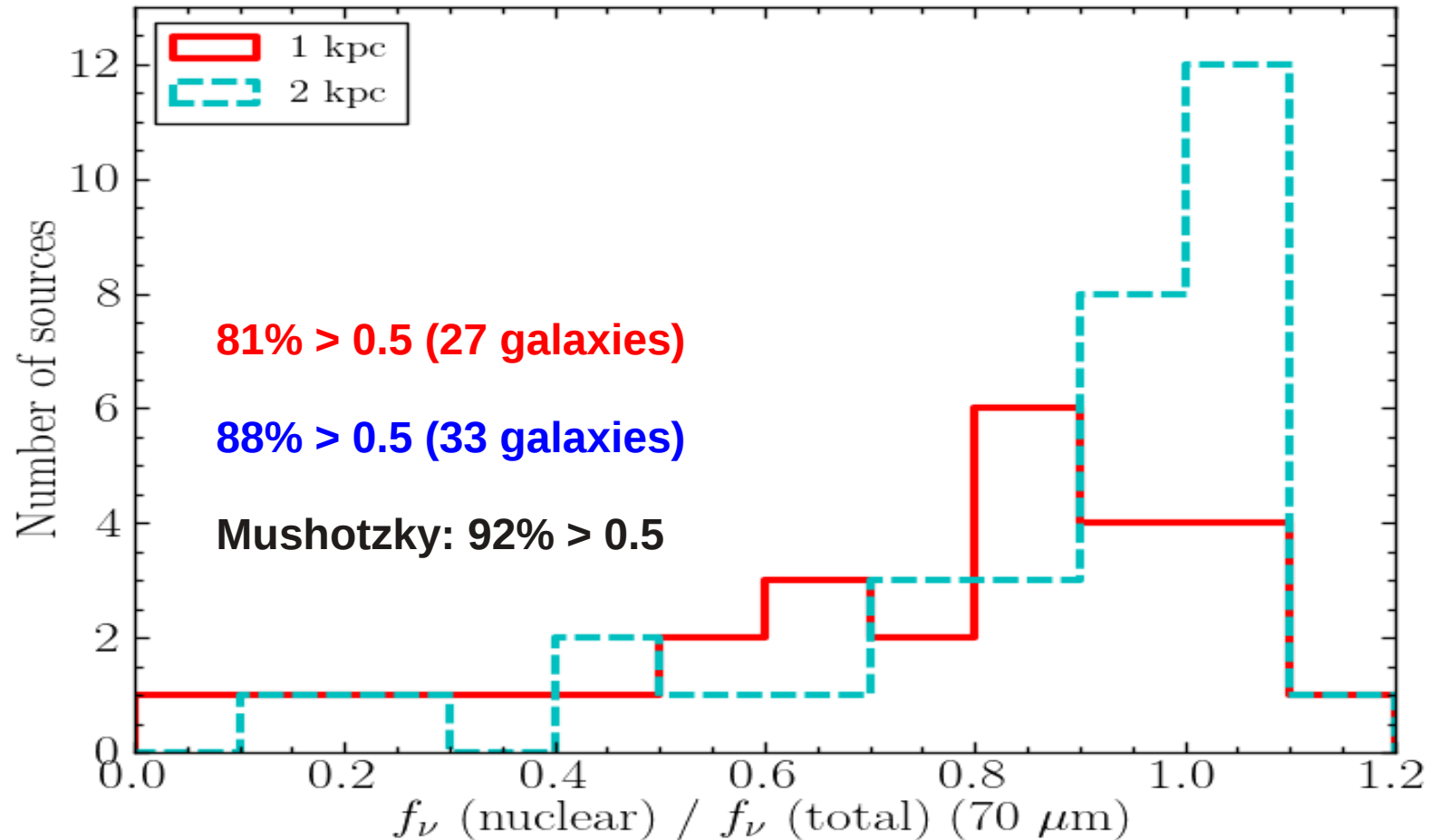
Angular resolution of 5.6 arcsec at 70 μm , median physical resolution of 0.8 kpc.

The sample and photometry



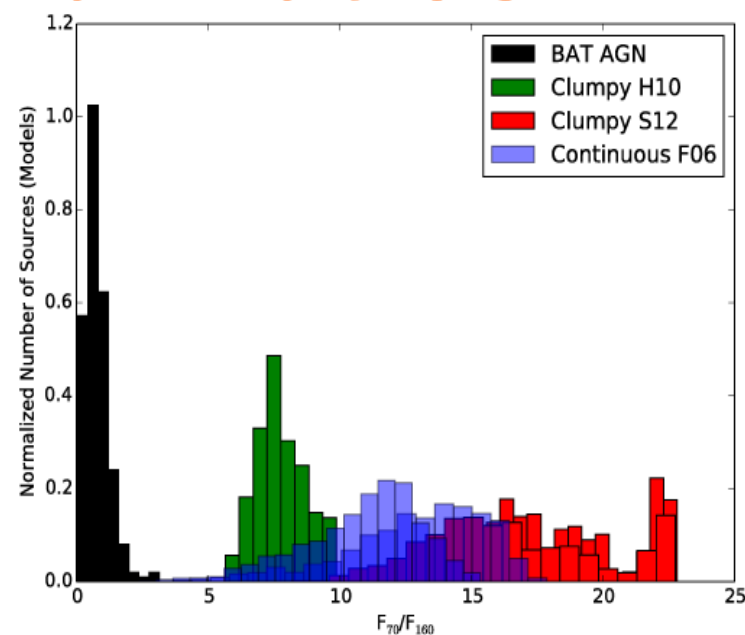
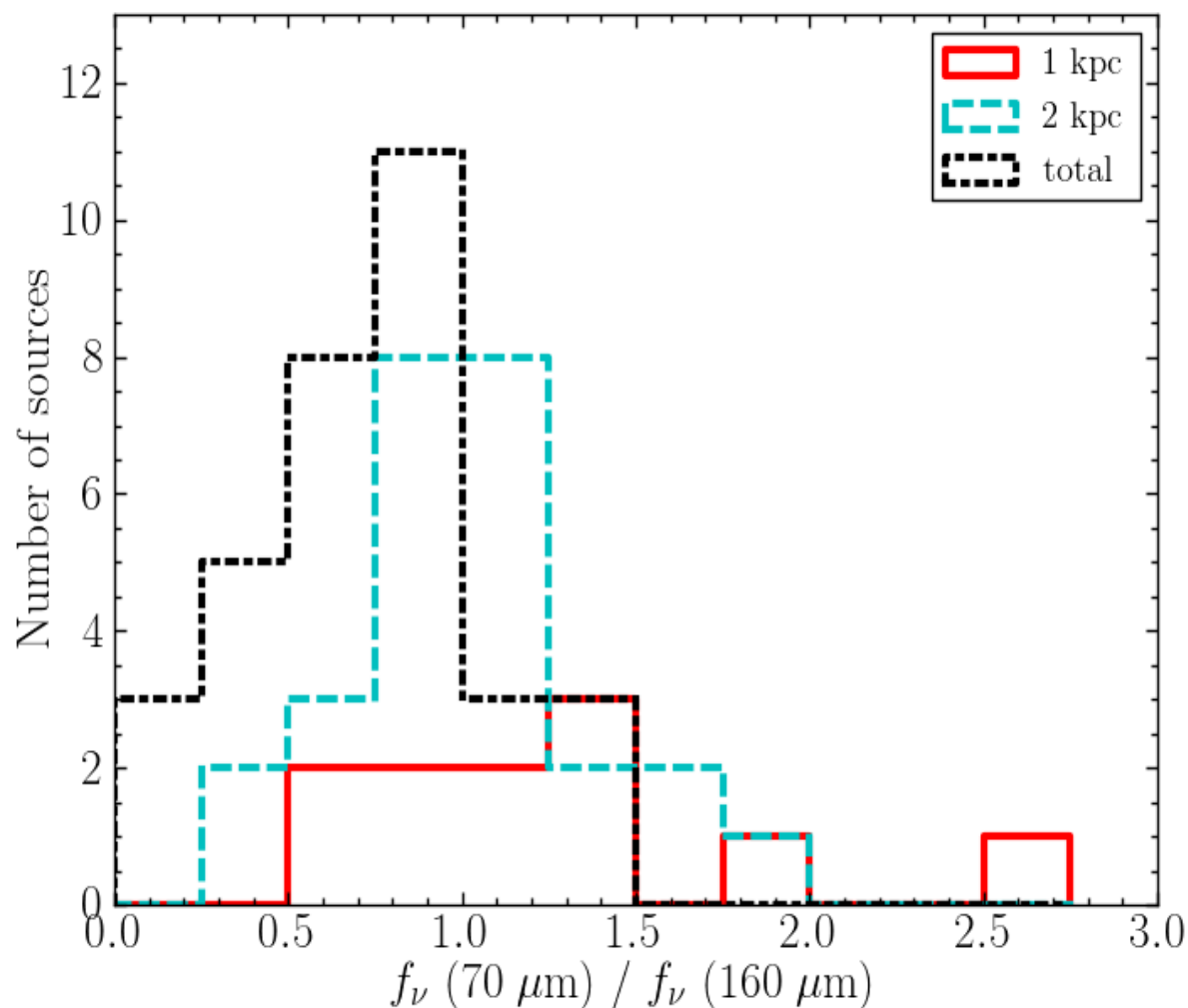
Aperture photometry for a radius of 1 kpc, 2 kpc, and the total galaxy (apertures with a diameter higher than 1.5 times the FWHM of each band)

Results: unresolved 70 micron emission

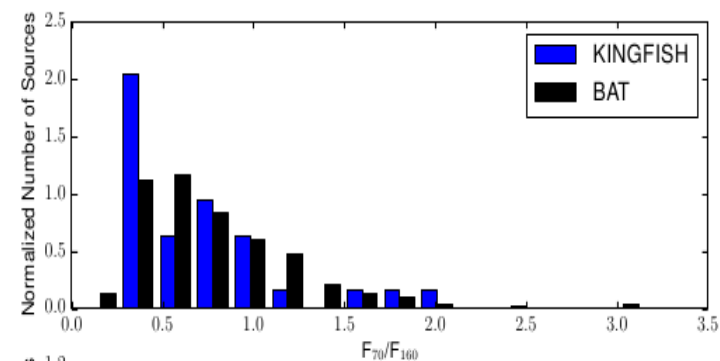


39 % are point-like at 70 microns (FWHM < 6") → the 70 μm emission originates in regions with sizes (diameters) of less than $\sim 0.5 - 2$ kpc

Results: 70/160 μm flux ratios

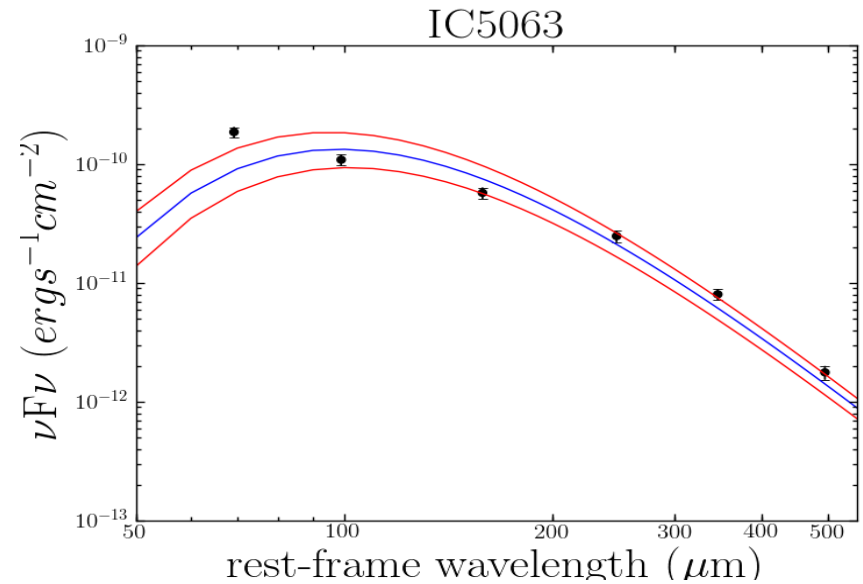
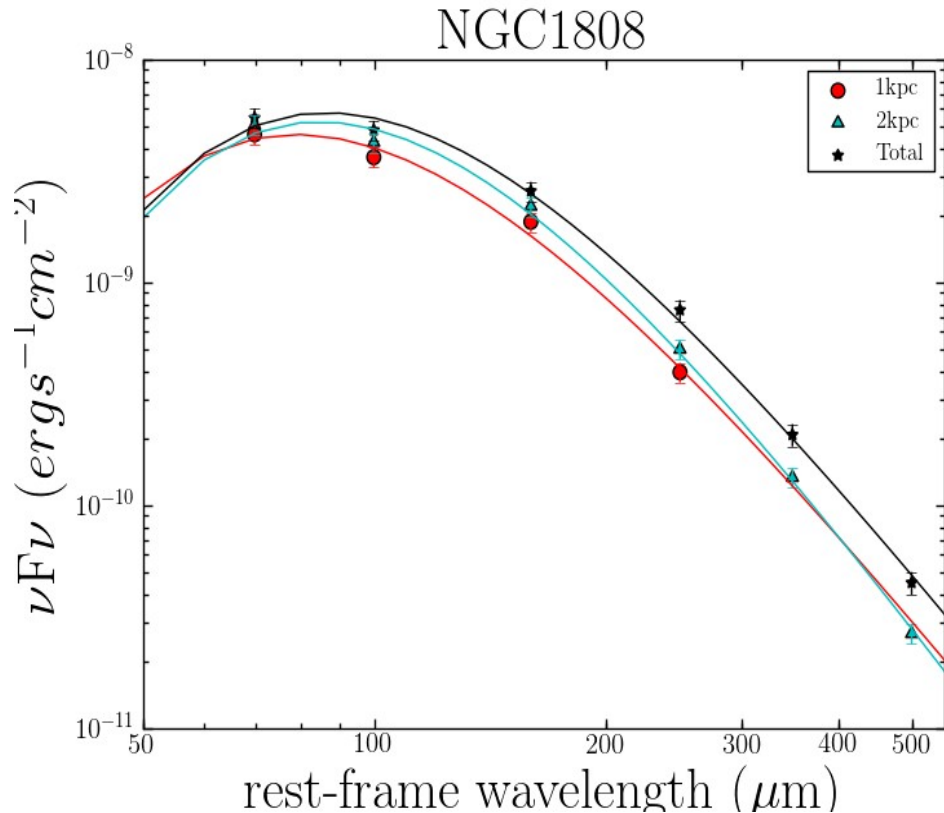


Meléndez et al. (2014)

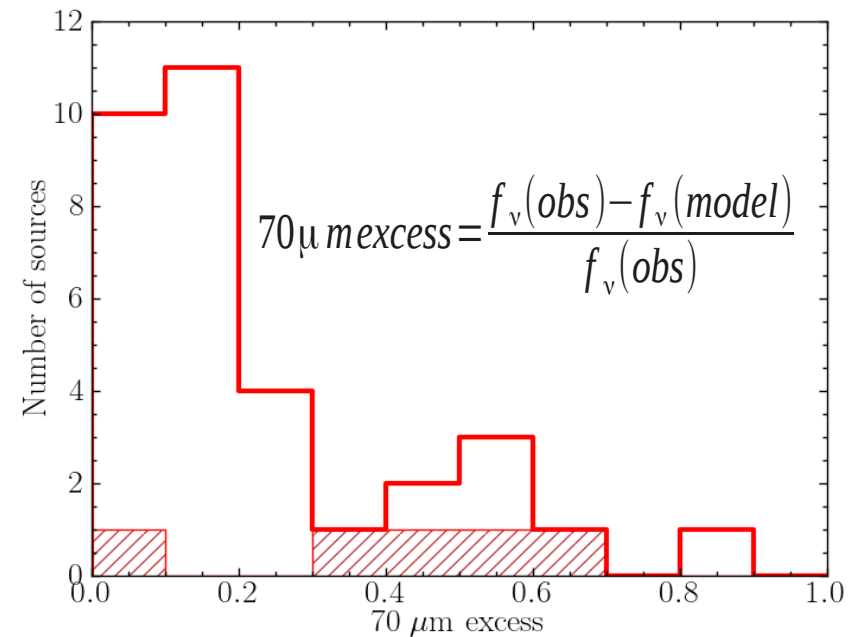


K-S test: 70/160 μm flux ratios of our galaxies, the BAT AGN sample and the normal galaxies are statistically indistinguishable.

Grey-body fitting



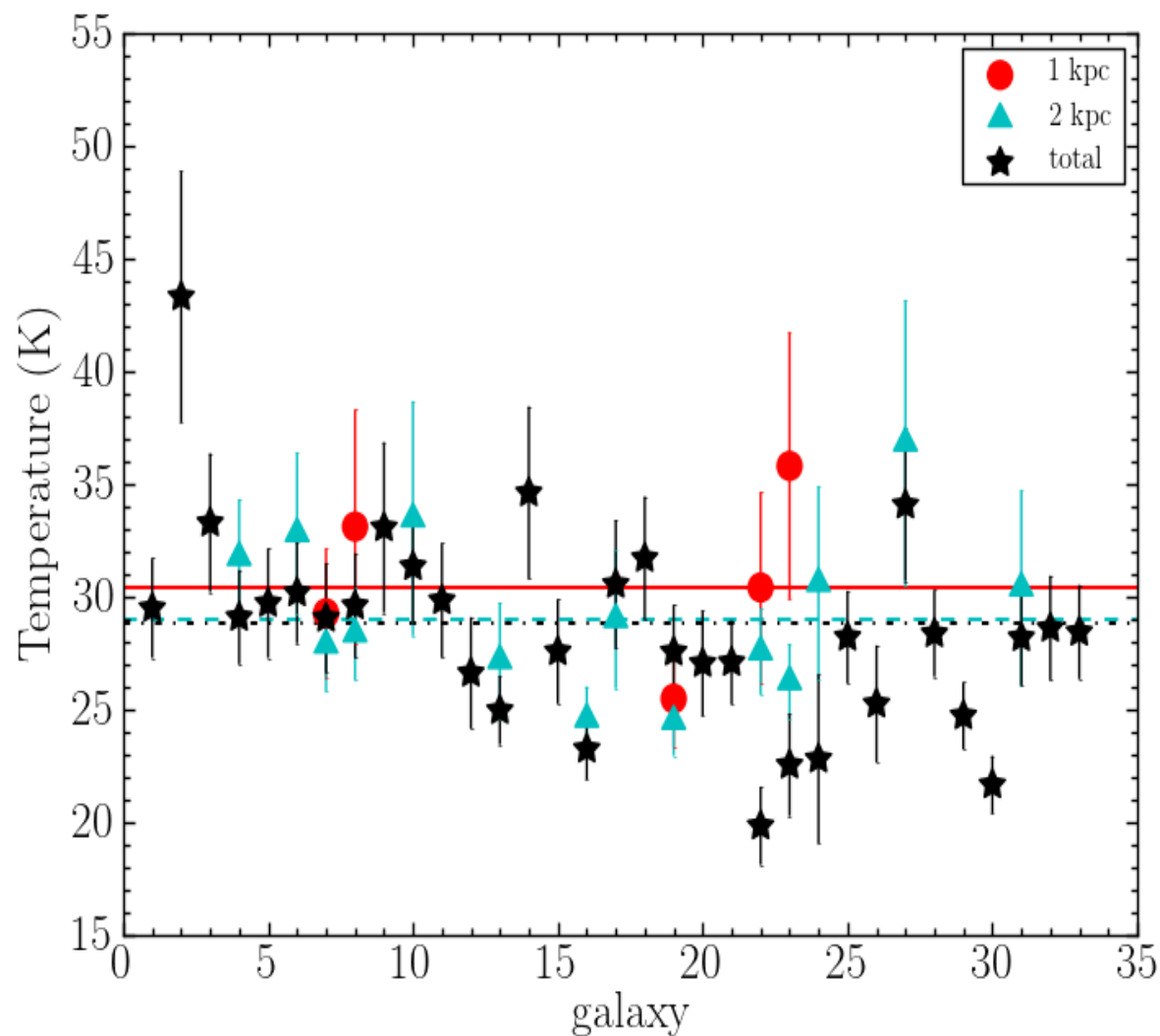
Grey-body with dust emissivity $\beta=2$



Fits are always better
without 70 μm data point

Evidence of AGN dust
heating *contamination* at
70 μm ?

Results: grey-body fitting



Free parameter β

Dust temperature decreases
from 1 kpc to 2 kpc and to
integrated

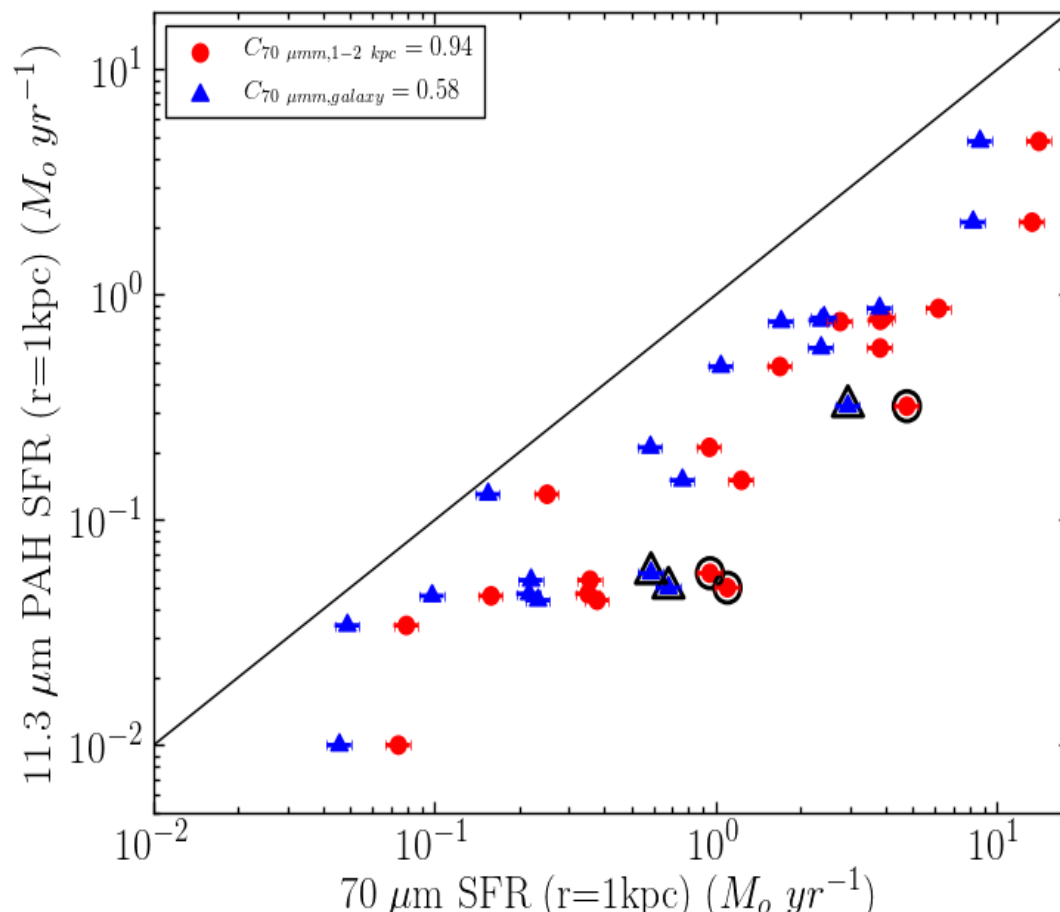
Median temperature:

1 kpc: 30 K

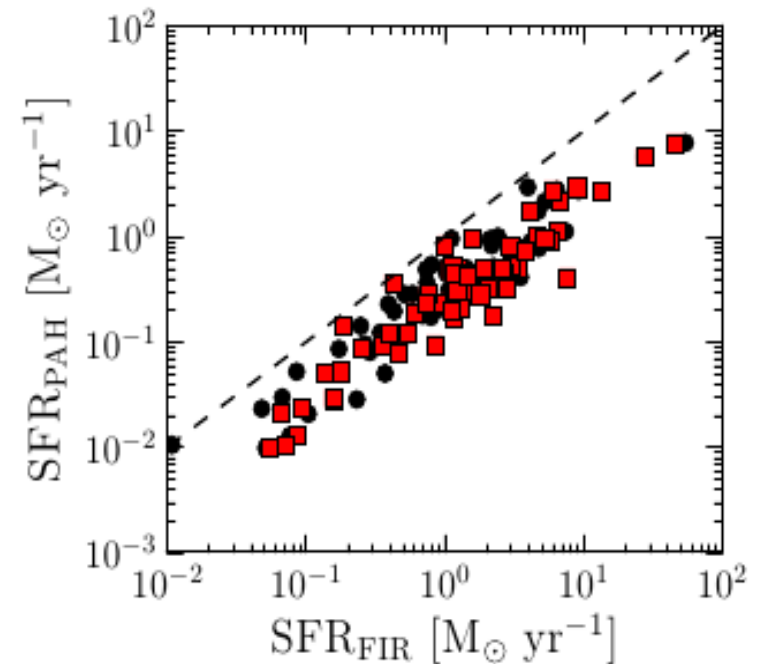
2 kpc: 29 K

Total : 28 K

Comparison of nuclear SFR



Muzotsky et al. (2014)



Discrepancy with the PAH derived SFR could be due to:

PAHs are being destroyed by the AGN.

There is a systematic in the calibrations.

There is an AGN contribution at 70 μm.

It is due to the aperture correction.

$$\text{SFR}(70 \mu\text{m})(M_{\odot} \text{ yr}^{-1}) = C_{70 \mu\text{m}, \text{region}} \times 10^{-43} \times L(70 \mu\text{m})(\text{erg s}^{-1})$$

AGN-dominated nuclear 70 μm emission galaxies

- Elevated 70/160 μm flux ratios
(8 galaxies)
- Nuclear dust temperature higher than typical star forming galaxies
(14 galaxies)
- 70 μm excess emission with respect to the fit of the FIR SEDs with a grey body with $\beta=2$
(7 galaxies)
- Comparison of nuclear SFR obtained from 70 μm and mid-IR indicators
(3 galaxies)

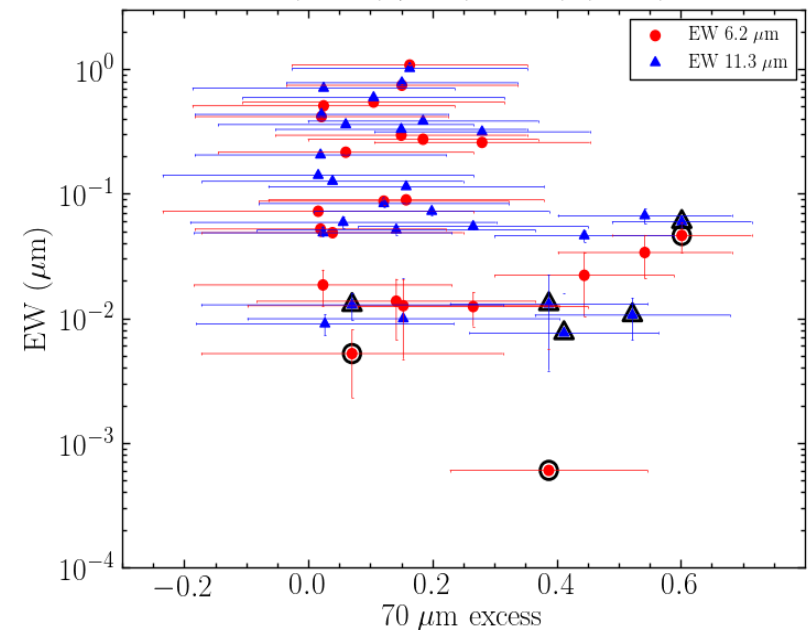
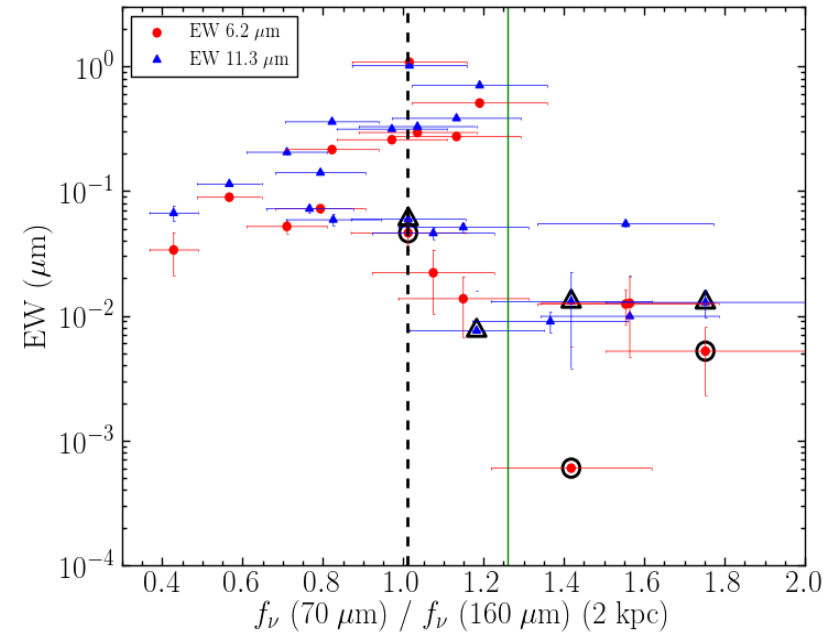


Table 7. Summary of criteria to select galaxies with a significant AGN contribution at $70\,\mu\text{m}$. In bold are marked galaxies satisfying at least half of the criteria.

Galaxy	#1 $f_\nu(70\,\mu\text{m})/f_\nu(160\,\mu\text{m})$			#2 $T_{\text{region}}/T_{\text{tot}}$		#3 $70\,\mu\text{m}$ excess	#4 $\text{SFR}_{70\,\mu\text{m}}/\text{SFR}_{\text{D-S}}$	criteria
	$r = 1\,\text{kpc}$	$r = 2\,\text{kpc}$	tot	$r = 1\,\text{kpc}$	$r = 2\,\text{kpc}$		$r = 1\,\text{kpc}$	
ESO323-G077	-	-	x	-	-	x	-	0/2
IC5063	-	-	✓	-	-	✓	-	2/2
Mrk1066	-	-	✓	-	-	x	-	1/2
NGC1068	✓	✓	x	-	x	x	-	1/3
NGC1320	-	x	x	-	-	x	-	0/2
NGC1365	x	x	x	-	x	x	x	0/4
NGC1386	x	x	x	x	x	x	x	0/4
NGC1808	x	x	x	x	x	x	-	0/3
NGC2110	-	x	x	-	-	x	-	0/2
NGC2273	-	x	x	-	x	x	x	0/4
NGC2992	-	x	x	-	-	x	x	0/3
NGC3081	-	x	x	-	-	✓	x	1/3
NGC3227	x	x	x	-	x	x	x	0/4
NGC3281	-	✓	✓	-	-	x	x	1/3
NGC3783	-	✓	x	-	-	x	✓	2/3
NGC4051	x	x	x	-	x	x	x	0/4
NGC4151	✓	✓	x	-	x	x	✓	2/4
NGC4253	-	-	✓	-	-	x	-	1/2
NGC4258	x	x	x	x	x	✓	x	1/4
NGC4388	-	x	x	-	-	x	x	0/3
NGC4507	-	-	x	-	-	x	-	0/2
NGC4579	x	x	x	✓	✓	✓	x	2/4
NGC4594	x	x	x	✓	x	x	x	1/4
NGC4725	-	x	x	-	✓	✓	x	2/4
NGC5135	-	-	x	-	-	x	-	0/2
NGC5347	-	x	x	-	-	✓	-	1/2
NGC5506	-	✓	✓	-	x	x	x	1/4
NGC7130	-	-	x	-	-	x	-	0/2
NGC7172	-	x	x	-	-	x	x	0/3
NGC7213	-	x	x	-	-	x	x	0/3
NGC7465	-	x	x	-	x	x	-	0/3
NGC7479	-	x	x	-	-	✓	✓	2/3
NGC7582	x	x	x	-	-	x	x	0/3

Table 7. Summary of criteria to select galaxies with a significant AGN contribution at $70\mu\text{m}$. In bold are marked galaxies satisfying at least half of the criteria.

Galaxy	#1 $f_\nu(70\mu\text{m})/f_\nu(160\mu\text{m})$			#2 $T_{\text{region}}/T_{\text{tot}}$		#3 $70\mu\text{m}$ excess	#4 $\text{SFR}_{70\mu\text{m}}/\text{SFR}_{\text{D-S}}$ $r = 1\text{ kpc}$	criteria
	$r = 1\text{ kpc}$	$r = 2\text{ kpc}$	tot	$r = 1\text{ kpc}$	$r = 2\text{ kpc}$			
ESO323-G077	-	-	x	-	-	x	-	0/2
IC5063	-	-	✓	-	-	✓	-	2/2
Mrk1066	High nuclear SFR		✓	-	-	x	-	1/2
NGC1068			x	-	x	x	-	1/3
NGC1320	-	x	x	-	-	x	-	0/2
NGC1365	x	x	x	-	x	x	x	0/4
NGC1386	x	x	x	x	x	x	x	0/4
NGC1808	x	x	x	x	x	x	-	0/3
NGC2110	-	x	x	-	-	x	-	0/2
NGC2273	-	x	x	-	x	x	x	0/4
NGC2992	-	x	x	-	-	x	x	0/3
NGC3081	-	x	x	-	-	✓	x	1/3
NGC3227	x	x	x	-	x	x	x	0/4
NGC3281	-	✓	✓	-	-	x	x	1/3
NGC3783	-	✓	x	-	-	x	✓	2/3
NGC4051	x	x	x	-	x	x	x	0/4
NGC4151	✓	✓	x	-	x	x	✓	2/4
NGC4253	High nuclear SFR		✓	-	-	x	-	1/2
NGC4258			x	x	x	✓	x	1/4
NGC4388	-	x	x	-	-	x	x	0/3
NGC4507	-	-	-	-	-	x	-	0/2
NGC4579	70 μm excess > nuclear flux		-	-	✓	✓	x	2/4
NGC4594			-	-	x	x	x	1/4
NGC4725	70 μm excess > nuclear flux		-	-	✓	✓	x	2/4
NGC5135			x	-	-	x	-	0/2
NGC5347	-	x	x	-	-	✓	-	1/2
NGC5506	-	✓	✓	-	x	x	x	1/4
NGC7130	-	-	x	-	-	x	-	0/2
NGC7172	-	x	x	-	-	x	x	0/3
NGC7213	-	x	x	-	-	x	x	0/3
NGC7465	-	x	x	-	x	x	-	0/3
NGC7479	-	x	x	-	-	✓	✓	2/3
NGC7582	x	x	x	-	-	x	x	0/3

Candidates to nuclear 70 μm AGN-dominated emission

- 15 galaxies (45% of the sample) satisfy at least one of these criteria
- 9 of them satisfy half of the criteria
- 5 RSA Seyfert galaxies (15% of the initial sample) whose nuclear ($r = 1\text{--}2$ kpc) 70 μm emission has a significant ($\sim 40\text{--}50\%$) contribution from hotter dust, presumably heated by the AGN:
 - IC 5063
 - NGC 3783
 - NGC 4151
 - NGC 5347
 - NGC 7479

None of them show 11.3 μm PAH emission on scales of less than 100 pc or high nuclear SFR.

Summary

- We have studied the nuclear and integrated 70-500 μm emission of 33 local Seyfert galaxies.
- 88% of the sample have a nuclear ($r = 2$ kpc) contribution to the total 70 μm emission greater than 50%.
- The majority of the galaxies have 70/160 μm flux ratios and dust temperatures ($20 < T < 43$ K) compatible with those of normal galaxies.
- We put forward four criteria to select galaxies with nuclear 70 μm AGN dominated emission: (1) elevated 70/160 μm flux ratios, (2) nuclear dust temperature higher than typical star forming galaxies, (3) 70 μm excess emission with respect to the fit of the FIR SEDs with a grey body with $\beta=2$, and (4) comparison of nuclear SFR obtained from 70 μm and mid-IR indicators.
- 15 galaxies (45% of the sample) satisfy at least one of these criteria, whereas 9 satisfy half or more of the criteria.
- We select 5 bona-fide candidates: IC 5063, NGC 3783, NGC 4151, NGC 5347, and NGC 7479.